



Повышение разрешения поточковых нейросетевых моделей для обработки данных РСДБ наблюдений проекта Радиоастрон

Кац Лев

Задача восстановления изображения

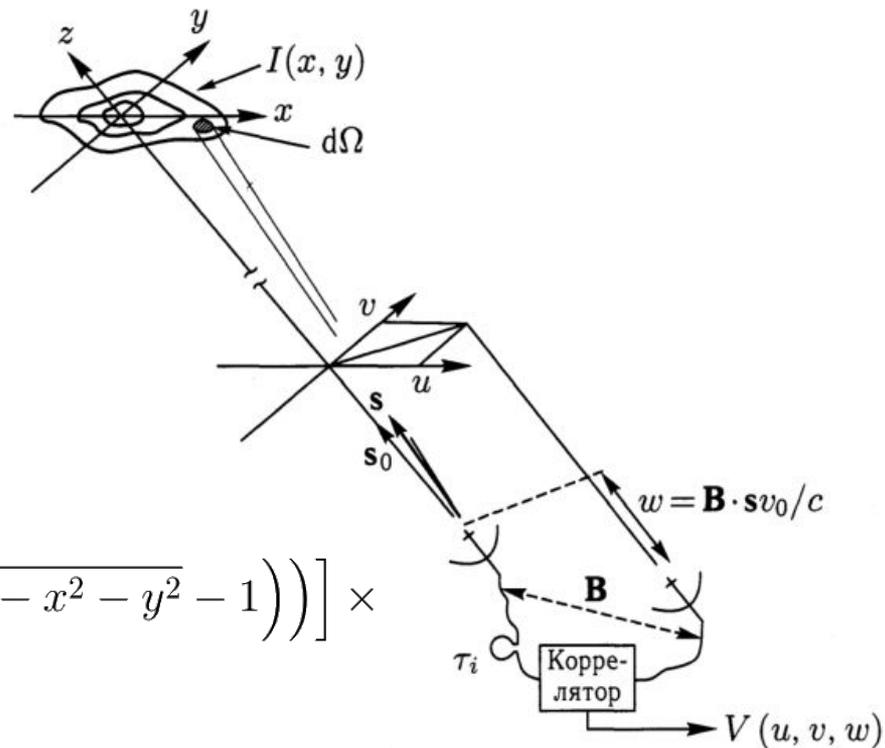


$$y = f(x)$$

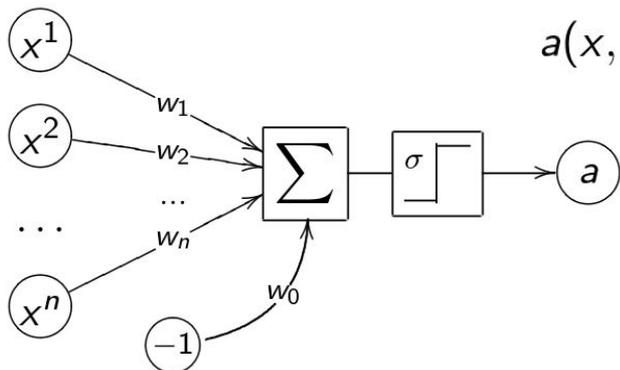
$$V(u, v, w) = \iint_{x^2+y^2 < 1} A(x, y) I(x, y) \times$$

$$\times \exp \left[-2\pi i \left(ux + vy + w \left(\sqrt{1 - x^2 - y^2} - 1 \right) \right) \right] \times$$

$$\times \frac{dx dy}{\sqrt{1 - x^2 - y^2}}$$



Модель нейрона.

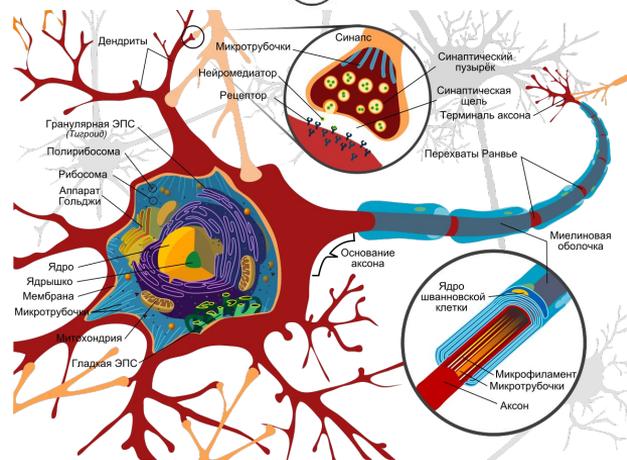
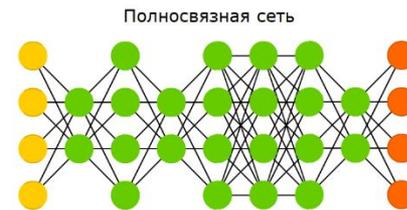


$$a(x, w) = \sigma(\langle w, x \rangle) = \sigma\left(\sum_{j=1}^n w_j f_j(x) - w_0\right),$$

МакКаллок-Питтс, 1943

$\sigma(z)$ — функция активации. Она обязательно нелинейная;
 w_j — весовые коэффициенты;
 w_0 — порог активации нейрона.

- Входной слой
- Скрытый слой
- Выходной слой



Теорема Цыбенко (1989): Нейронная сеть с одним скрытым слоем может равномерно аппроксимировать любую непрерывную функцию на отрезке с заданной точностью

Постановка задачи. Апостериорное распределение

$$\theta^* = \arg \min_{\theta} D_{\text{KL}}(q_{\theta}(x) \| p(x|y))$$

$$= \arg \min_{\theta} \int q_{\theta}(x) [\log q_{\theta}(x) - \log p(x|y)] dx$$

...

$$x \sim q_{\theta}(x) \Leftrightarrow x = G_{\theta}(z), z \sim \mathcal{N}(0, 1)$$

...

$$= \arg \min_{\theta} \mathbb{E}_{z \sim \mathcal{N}(0,1)} \left\{ -\log p(y|G_{\theta}(z)) - \log p(G_{\theta}(z)) \right. \\ \left. + \log \pi(z) - \log \left| \det \frac{dG_{\theta}(z)}{dz} \right| \right\}.$$

Расстояние Кúльбака — Лéйблера

$$D_{\text{KL}}(P \| Q) = \int_X p(x) \log \frac{p(x)}{q(x)} dx,$$

, где **P**, **Q** — плотности
распределения

$$D_{\text{KL}}(P \| Q) \geq 0, :$$

$$D_{\text{KL}}(P \| Q) = 0 \iff P = Q$$

$$D_{\text{KL}}(P \| Q) \neq D_{\text{KL}}(Q \| P).$$

Регуляризации как плотность распределения

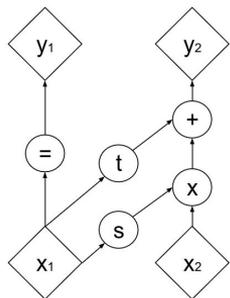
$$\theta^* = \arg \min_{\theta} \mathbb{E}_{z \sim \mathcal{N}(0,1)} \left\{ -\log p(y|G_{\theta}(z)) - \log p(G_{\theta}(z)) \right. \\ \left. + \log \pi(z) - \log \left| \det \frac{dG_{\theta}(z)}{dz} \right| \right\}$$

$$\theta^* = \arg \min_{\theta} \sum_{k=1}^N \left\{ \mathcal{L}(y, f(G_{\theta}(z_k))) + \lambda \mathcal{R}(G_{\theta}(z_k)) - \log \left| \det \frac{dG_{\theta}(z_k)}{dz_k} \right| \right\},$$

$$-\log p(G_{\theta}(z)) \quad \longrightarrow \quad \lambda \mathfrak{R}(G_{\theta}(z))$$

$$\mathfrak{R}(x) = \lambda_1 \text{ll_loss}(x) + \lambda_2 \text{loss_TSV}(x) + \lambda_3 \text{loss_center}(x) + \\ + \lambda_4 \text{loss_flux}(x) + \lambda_5 \text{loss_ce}(x)$$

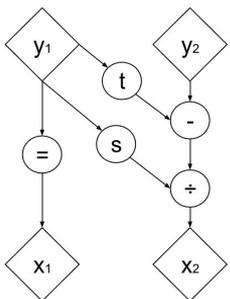
Affine Coupling layer и Real NVP



$$y_{1:d} = x_{1:d}$$

$$y_{d+1:D} = x_{d+1:D} \odot \exp(s(x_{1:d})) + t(x_{1:d}),$$

forward



inverse

$$\frac{\partial y}{\partial x^T} = \begin{bmatrix} \mathbb{I}_d & 0 \\ \frac{\partial y_{d+1:D}}{\partial x_{1:d}^T} & \text{diag}(\exp[s(x_{1:d})]) \end{bmatrix},$$

$$\det \frac{\partial y}{\partial x^T} = \exp \left[\sum_j s(x_{1:d})_j \right].$$

Diederik P. Kingma*, Prafulla Dhariwal*
OpenAI, San Francisco

Увеличение размеров генерируемых изображений

Abstract

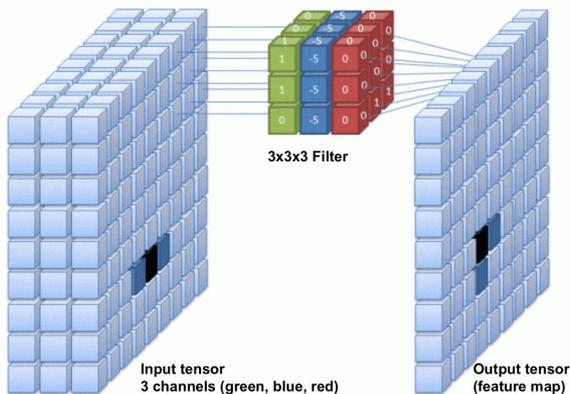
Flow-based generative models (Dinh et al., 2014) are conceptually a tractability of the exact log-likelihood, tractability of exact latent-var and parallelizability of both training and synthesis. In this paper we propose *Glow*, a simple type of generative flow using an invertible 1×1 method we demonstrate a significant improvement in lo benchmarks. Perhaps most strikingly, we demonstrate optimized towards the plain log-likelihood objective is ca looking synthesis and manipulation of large images. T available at <https://github.com/openai/glow>.

$$\log \left| \det \left(\frac{d \text{conv2D}(\mathbf{h}; \mathbf{W})}{d \mathbf{h}} \right) \right| = h \cdot w \cdot \log |\det(\mathbf{W})|$$

1 Introduction

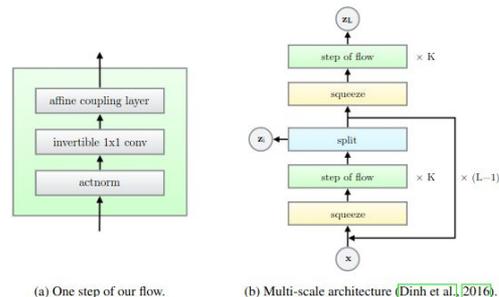
Two major unsolved problems in the field of machine learning are (1) learning from few datapoints, like humans; and (2) generalization: robustness to changes of the task or its context. AI systems, for example, often do not work at all when given inputs that are different from their training distribution. A promise of *generative models*, a major branch of machine learning,

*Equal contribution.



$$\mathbf{W} = \mathbf{P}\mathbf{L}(\mathbf{U} + \text{diag}(\mathbf{s}))$$

$$\log |\det(\mathbf{W})| = \text{sum}(\log |\mathbf{s}|)$$



Увеличения размеров генерируемых изображений

Wavelet Flow: Fast Training of High Resolution Normalizing Flows

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³Borealis AI, ⁴Samsung AI Centre Toronto, ⁵Vector Institute

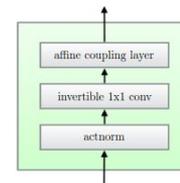
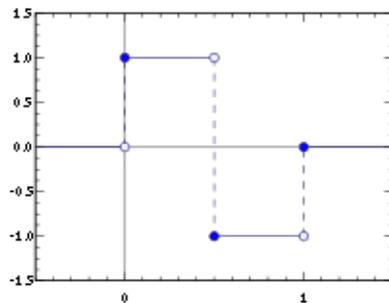
jyyu@eecs.yorku.ca kostas@cs.ryerson.ca mab@eecs.yorku.ca

Abstract

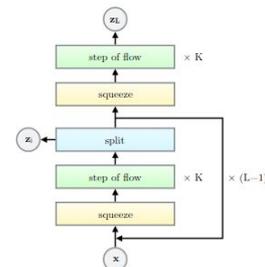
Normalizing flows are a class of probabilistic generative models which allow for both fast density computation and efficient sampling and are effective at modelling complex distributions like images. A drawback among current models is their significant training cost, sometimes requiring months of GPU training time to achieve state-of-the-art results. This paper introduces *Wavelet Flow*, a multi-scale, normalizing flow architecture based on wavelets. A Wavelet Flow has an explicit representation of signal scale that inherently includes models of lower resolution signals and conditional generation of higher resolution signals, *i.e.*, super resolution. A major advantage of Wavelet Flow is the ability to construct generative models for high resolution data (*e.g.*, 1024×1024 images) that are impractical with previous models. Furthermore, Wavelet Flow is competitive with previous normalizing flows in terms of bits per dimension on standard (low resolution) benchmarks while being up to $15\times$ faster to train.

1 Introduction

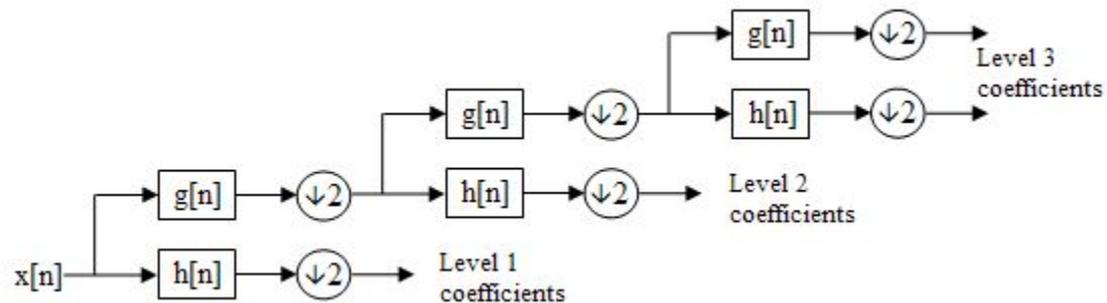
Here we introduce *Wavelet Flow*, a multi-scale, conditional normalizing flow architecture based on



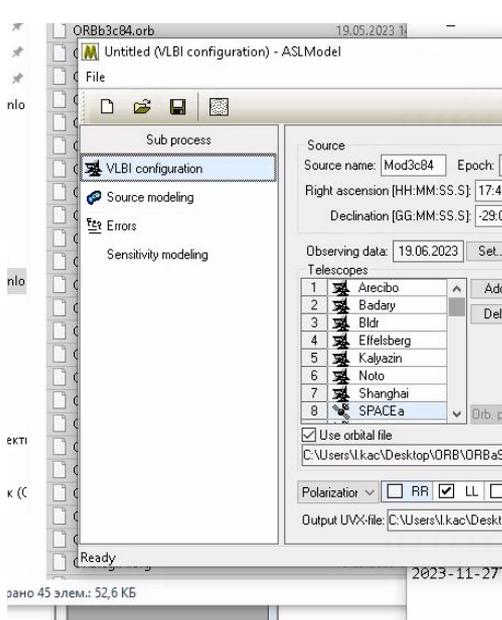
(a) One step of our flow.



(b) Multi-scale architecture (Dinh et al., 2016).



Создание модельных датасетов.



ORBB3c84.orb 19.05.2023 11

Untitled (VLBI configuration) - ASLModel

File

Sub process

- VLBI configuration
- Source modeling
- Errors
- Sensitivity modeling

Source

Source name: Mod3c84 Epoch: 20

Right ascension [HH:MM:SS.S]: 17.45.4

Declination [GG:MM:SS.S]: -23.00.2

Observing data: 19.06.2023 Set...

Telescopes

		Add...	Delete
1	Arecibo		
2	Badary		
3	Bldr		
4	Effelsberg		
5	Kalyazin		
6	Noto		
7	Shanghai		
8	SPACEa		

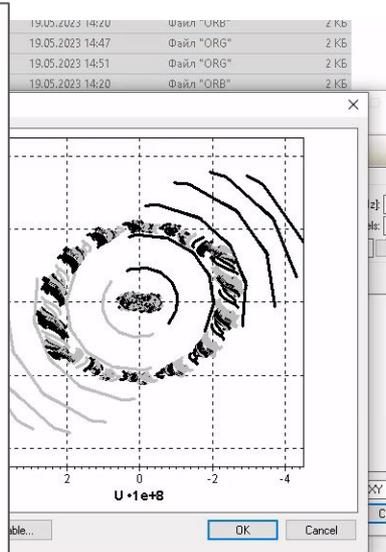
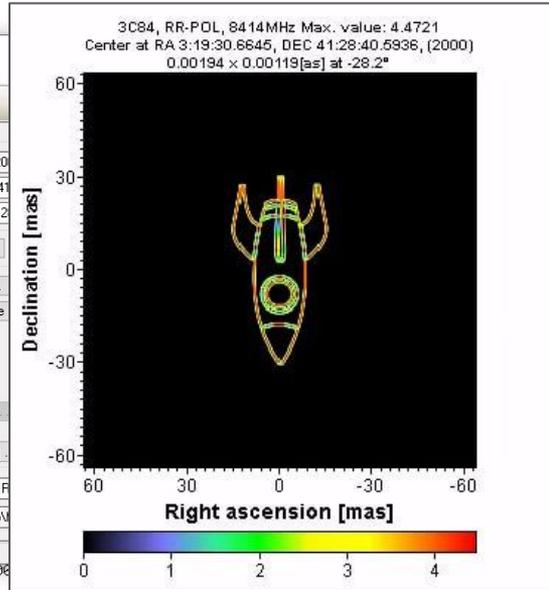
Use orbital file
C:\Users\N.kac\Desktop\ORB\ORBBaS...

Polarization RR LL P

Output UVX-file: C:\Users\N.kac\Desktop\...

Ready 2023-11-27T00:00:00

рано 45 элем.: 52,6 КБ



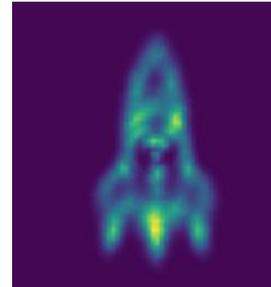
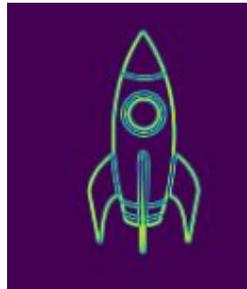
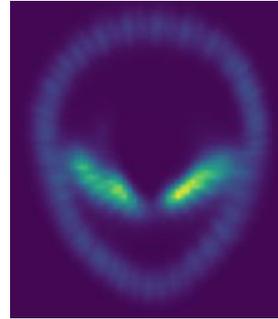
ORBB3c84.orb - Блокнот

Файл Правка Формат Вид Справка

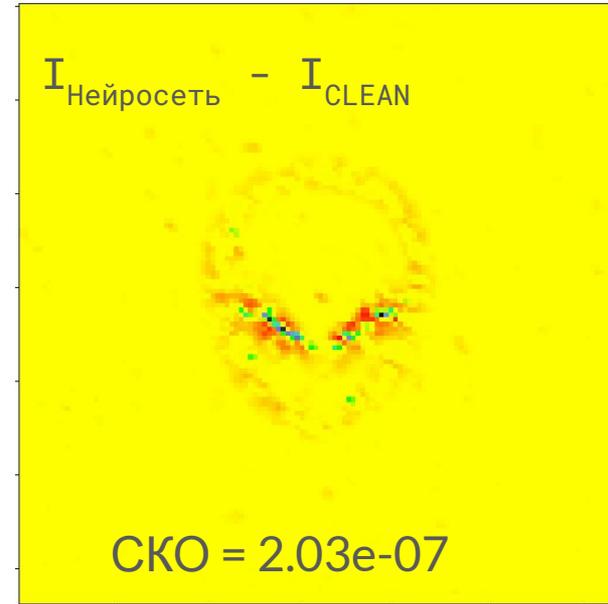
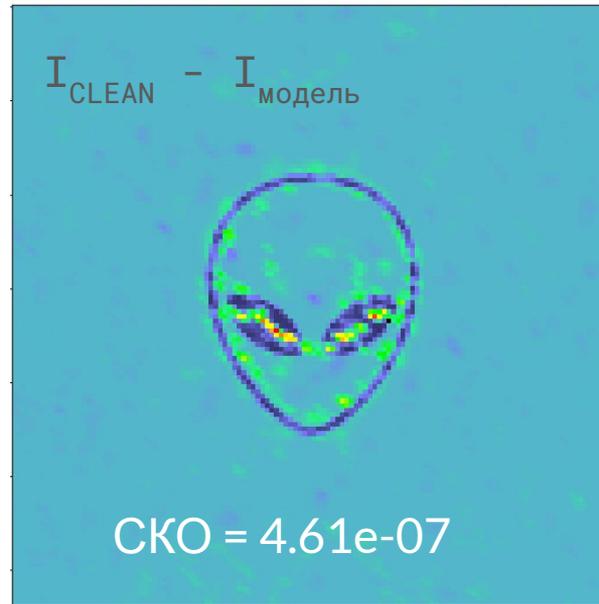
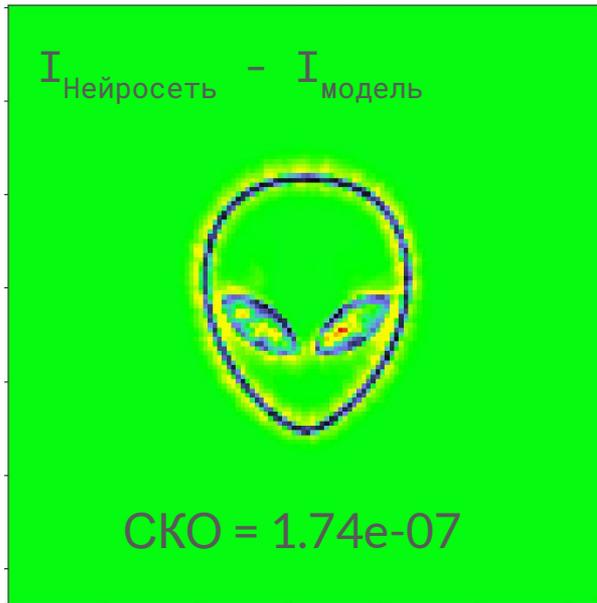
META_STOP

2023-11-01T00:00:00	8287.55	43706.8	-9554.31
2023-11-02T00:00:00	-5114.18	44865.8	5581.36
2023-11-03T00:00:00	-17503	37138.7	19611.6
2023-11-04T00:00:00	-26425.1	22055.8	29757.5
2023-11-05T00:00:00	-30113.4	2604.41	34009.5
2023-11-06T00:00:00	-27837.4	-17362.8	31525.6
2023-11-07T00:00:00	-20047.8	-33891	22797.6
2023-11-08T00:00:00	-8287.55	-43706.8	9554.31
2023-11-09T00:00:00	5114.18	-44865.8	-5581.36
2023-11-10T00:00:00	17503	-37138.7	-19611.6
2023-11-11T00:00:00	26425.1	-22055.8	-29757.5
2023-11-12T00:00:00	30113.4	-2604.41	-34009.5
2023-11-13T00:00:00	27837.4	17362.8	-31525.6
2023-11-14T00:00:00	20047.8	33891	-22797.6
2023-11-15T00:00:00	8287.55	43706.8	-9554.31
2023-11-16T00:00:00	-5114.18	44865.8	5581.36
2023-11-17T00:00:00	-17503	37138.7	19611.6
2023-11-18T00:00:00	-26425.1	22055.8	29757.5
2023-11-19T00:00:00	-30113.4	2604.41	34009.5
2023-11-20T00:00:00	-27837.4	-17362.8	31525.6
2023-11-21T00:00:00	-20047.8	-33891	22797.6
2023-11-22T00:00:00	-8287.55	-43706.8	9554.31
2023-11-23T00:00:00	5114.18	-44865.8	-5581.36
2023-11-24T00:00:00	17503	-37138.7	-19611.6
2023-11-25T00:00:00	26425.1	-22055.8	-29757.5
2023-11-26T00:00:00	30113.4	-2604.41	-34009.5
2023-11-27T00:00:00	27837.4	17362.8	-31525.6

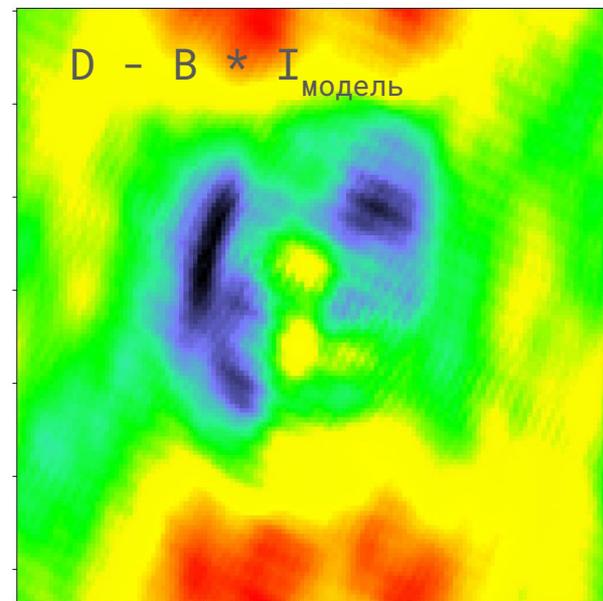
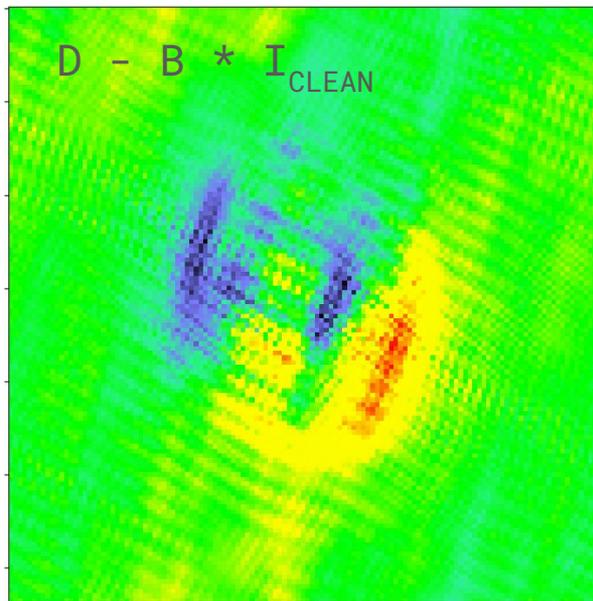
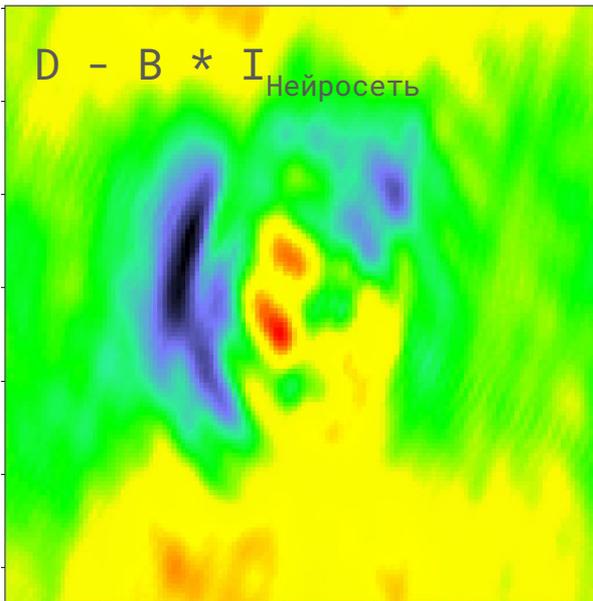
Восстановление модельных изображений



Сравнение с классическими методами

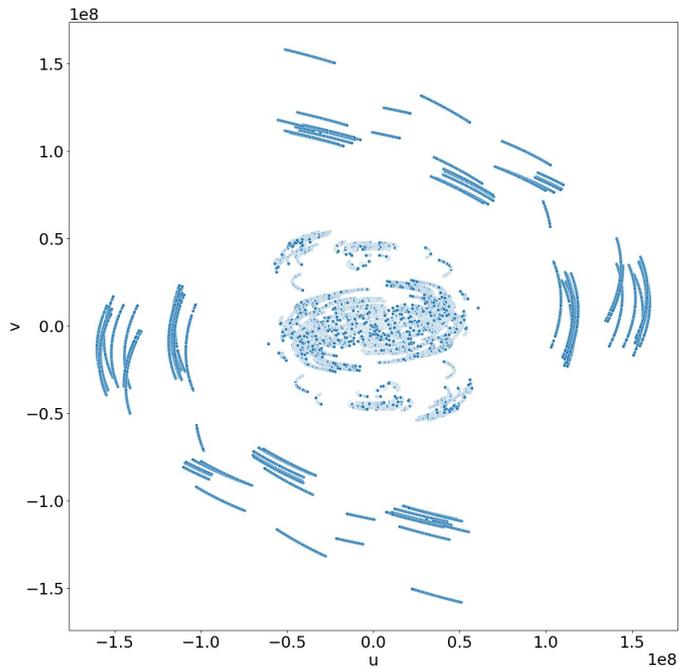


Сравнение с классическими методами

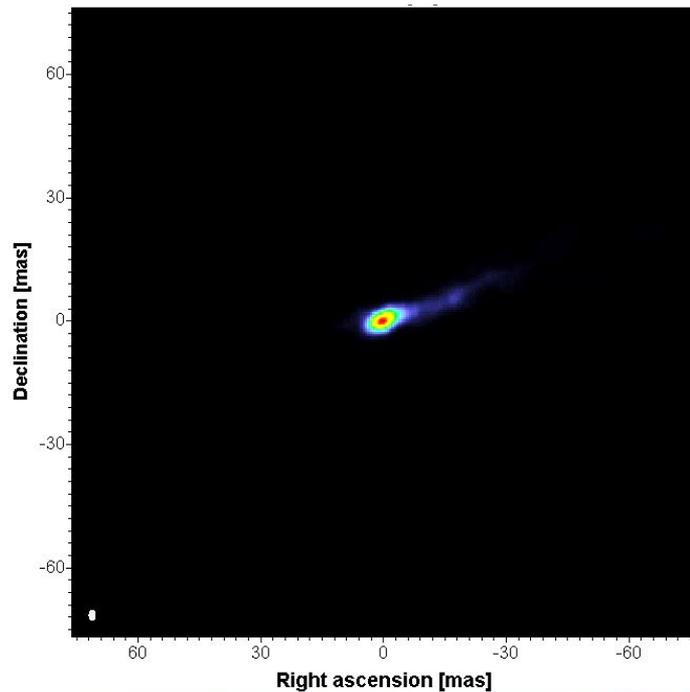


Данные наблюдений M87.

1228+126 LL-POL
1652.25 MHz
RA 12:30:49.4 (2000)
DEC 12:23:28.4 (2000)



CLEAN



Данные наблюдений. Шумы

$$\tilde{V}_{ik}(t) = g_i(t)g_k^*(t)V_{ik} + \varepsilon_{ik}(t).$$

\tilde{V}_{ik} — наблюдаемая видность
 V_{ik} — истинная видность

text file. For a sample Tsys format given below.

321 20:02:07 33.5 36.7 34.2 32.1 38.9 39.7 23.7 !

the index keywords could be used as follows:

INDEX = 'R1:8:2', 'R3', 'X', 'X', 'X', 'X', 'X'
 INDEX2 = 'L1:8:2', 'L3', 'X', 'X', 'X', 'L7', 'X'

This would map column 1 to IF 1-8 RCP and LCP, column 2 to IF 3 LCP and RCP, and column 6 to IF 7 LCP. The INDEX2 keyword is not necessary in all cases.

1) The CONTROL group

This must be the first group in the text file and is used to specify the default input format for the TSY5 and TANT entries which may follow. The CONTROL group is optional; if not specified it will be assumed that Tsys values are supplied for all IF's and polarizations present in the uv-data file in IF-polarization column order at each Tsys or Tant time stamp (ie. 'R1', 'L1', 'R2', 'L2', .. or 'R1', 'R2', depending on the number of polarizations). An alternative default format may be specified using keywords INDEX and INDEX2 in the control group, in the manner described above. This will then be used for all antennas for which a specific INDEX format is not given.

Example:
 CONTROL
 INDEX = 'R1', 'L1', 'R2', 'L2', 'R3', 'L3' /

Note: INDEX2 keyword can optionally be added but the CONTROL keyword is required. No other keywords are expected in the control group. Unrecognized entries in the input calibration file are simply counted and reported by number. This minimises the editing necessary in modifying VBA calibration files which may contain other information.

2) The BASELINE group

The baseline group is used to specify baseline-dependent amplitude gain corrections. These are entered directly into the BL table and

Тепловой шум:

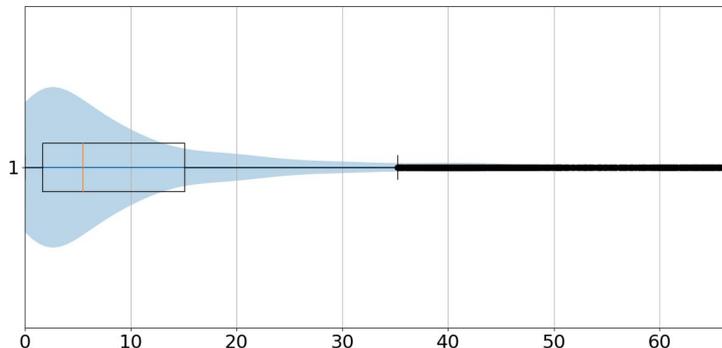
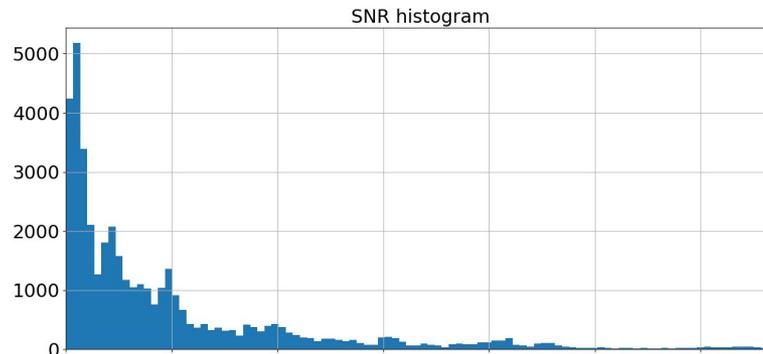
$$\sigma_{ij} = \frac{1}{\eta} \sqrt{\frac{\text{SEFD}_i \times \text{SEFD}_j}{2\Delta\nu \Delta t}}$$

SEFD — system equivalent flux density;

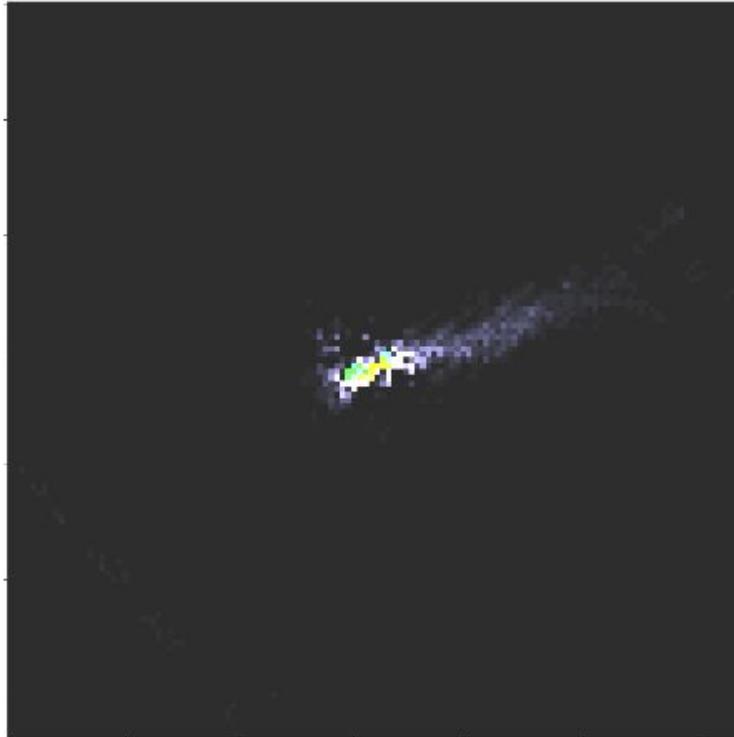
$\Delta\nu$ — ширина полосы;

Δt — интервал интегрирования;

$\eta = 0.88$.



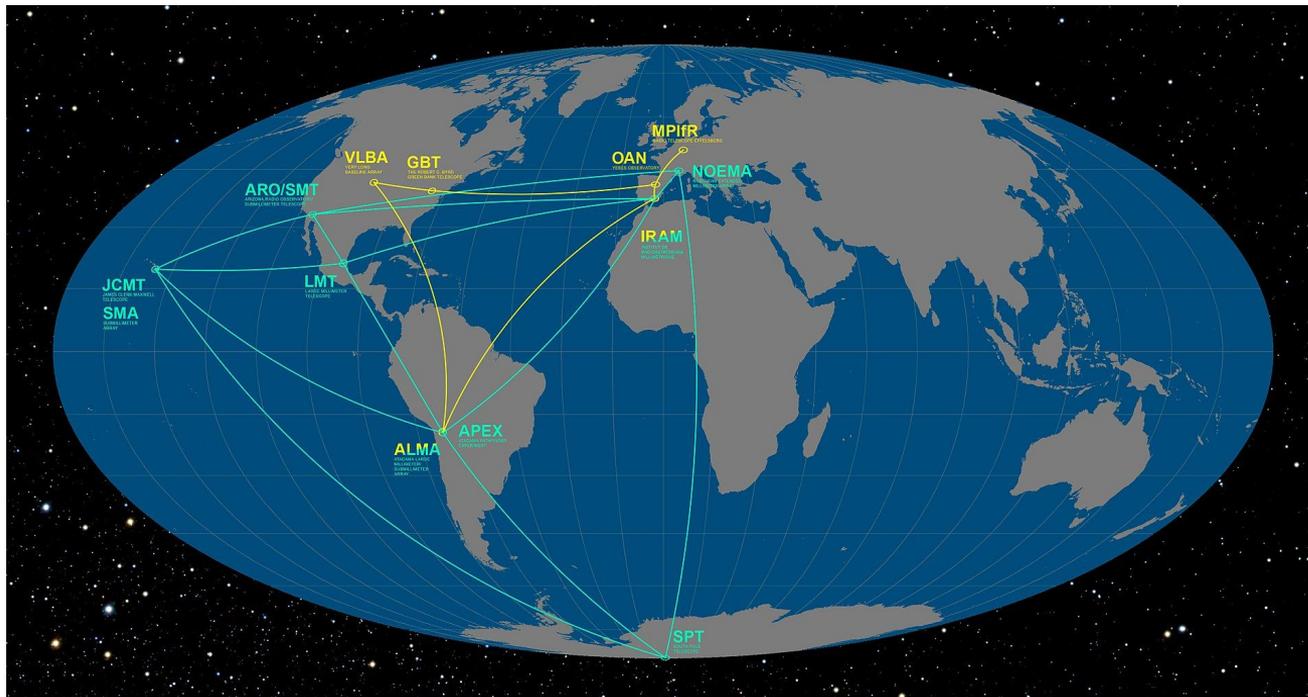
Данные наблюдений M87.



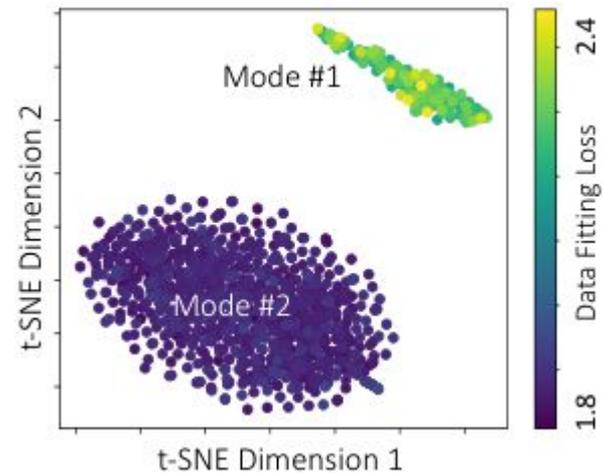
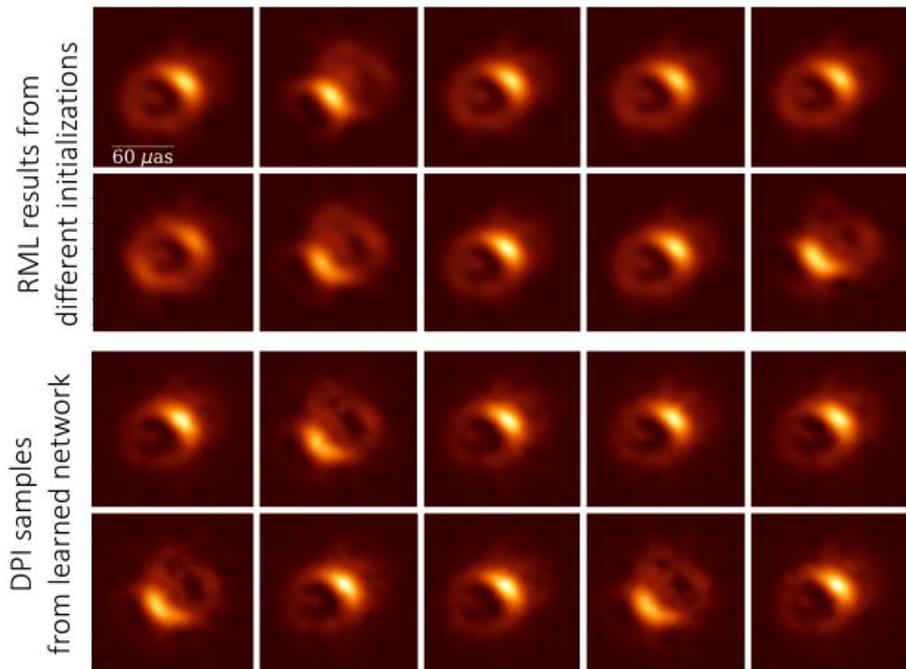
Планы

1. Дальнейшее **повышение разрешения** (сейчас **256x256**), вероятно потребует распределенного обучения
2. Подробное **исследование распределения изображений**, а не отдельных реализаций
3. **Учет систематических мультипликативных погрешностей** из-за неточной калибровки (обычно для этого используют **соотношения замыканий**)

Спасибо за внимание!

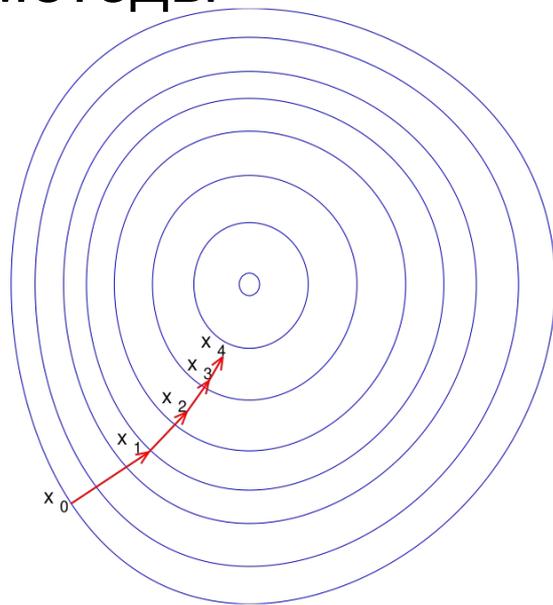


Достоверность деталей изображения



Процесс обучения. Градиентные методы

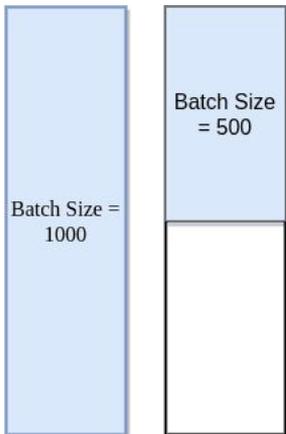
$$\begin{aligned} L(\mathbf{W}) &= \sum_{i=1}^N l(\mathbf{y}_i, f(\mathbf{W}, \mathbf{x}_i)) = \\ &= \{\text{например, если задача регрессии}\} = \\ &= \sum_{i=1}^N |y_i - f(\mathbf{W}, \mathbf{x}_i)|^2 \longrightarrow \min \end{aligned}$$



$$\mathbf{W}_k = \mathbf{W}_{k-1} - \gamma \nabla L(\mathbf{W}) \Big|_{\mathbf{W}_{k-1}}$$

Процесс обучения. Стохастический градиент

$$\begin{aligned} \frac{\partial}{\partial W} \mathbb{E} [l(\mathbf{y}_i, f(W, \mathbf{x}_i))] &= \frac{\partial}{\partial W} \mathbb{E} \left[\frac{1}{B} \sum_{i=1}^B l(\mathbf{y}_i, f(W, \mathbf{x}_i)) \right] = \\ &= \mathbb{E} \left[\frac{1}{B} \sum_{i=1}^B \frac{\partial}{\partial W} l(\mathbf{y}_i, f(W, \mathbf{x}_i)) \right] \approx \\ &\approx \frac{1}{B} \sum_{i=1}^B \frac{\partial}{\partial W} l(\mathbf{y}_i, f(W, \mathbf{x}_i)) \end{aligned}$$



Iterations per Epoch = 1

Iterations per Epoch = 2

...



Iterations per Epoch = 10

input : γ (lr), β_1, β_2 (betas), θ_0 (params), $f(\theta)$ (objective), ϵ (epsilon)
 λ (weight decay), *amsgrad*, *maximize*

initialize : $m_0 \leftarrow 0$ (first moment), $v_0 \leftarrow 0$ (second moment), $\widehat{v}_0^{max} \leftarrow 0$

for $t = 1$ to ... **do**

if *maximize* :

$g_t \leftarrow -\nabla_{\theta} f_t(\theta_{t-1})$

else

$g_t \leftarrow \nabla_{\theta} f_t(\theta_{t-1})$

$\theta_t \leftarrow \theta_{t-1} - \gamma \lambda \theta_{t-1}$

$m_t \leftarrow \beta_1 m_{t-1} + (1 - \beta_1) g_t$

$v_t \leftarrow \beta_2 v_{t-1} + (1 - \beta_2) g_t^2$

$\widehat{m}_t \leftarrow m_t / (1 - \beta_1^t)$

$\widehat{v}_t \leftarrow v_t / (1 - \beta_2^t)$

if *amsgrad*

$\widehat{v}_t^{max} \leftarrow \max(\widehat{v}_{t-1}^{max}, \widehat{v}_t)$

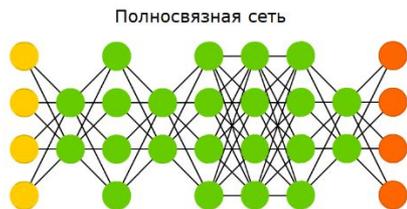
$\theta_t \leftarrow \theta_t - \gamma \widehat{m}_t / (\sqrt{\widehat{v}_t^{max}} + \epsilon)$

else

$\theta_t \leftarrow \theta_t - \gamma \widehat{m}_t / (\sqrt{\widehat{v}_t} + \epsilon)$

return θ_t

Объединение в слои

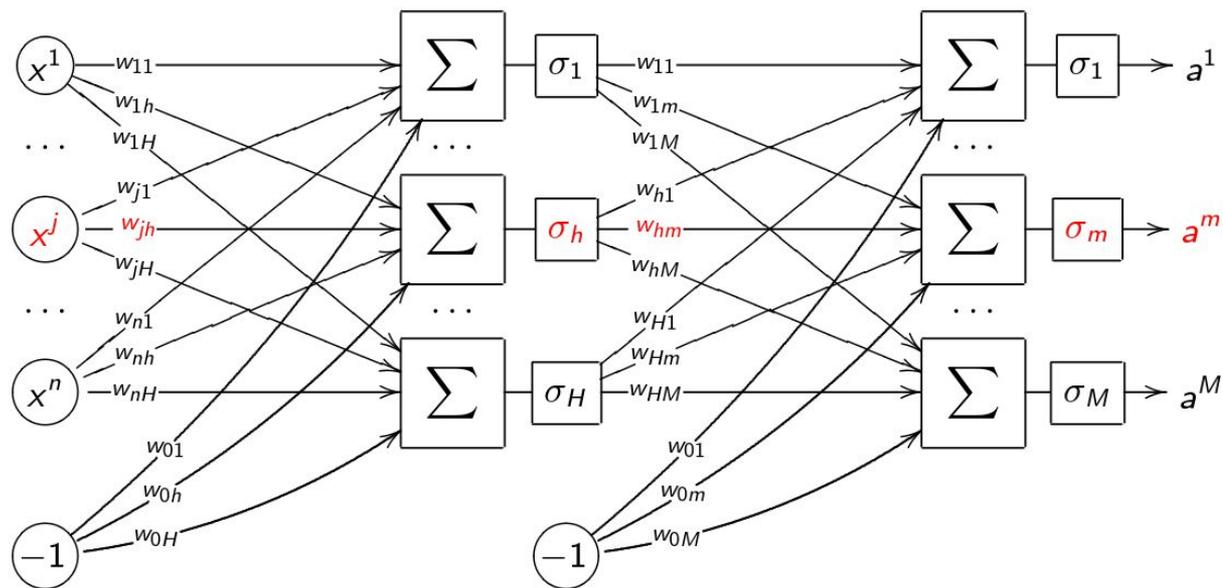


- Входной слой
- Скрытый слой
- Выходной слой

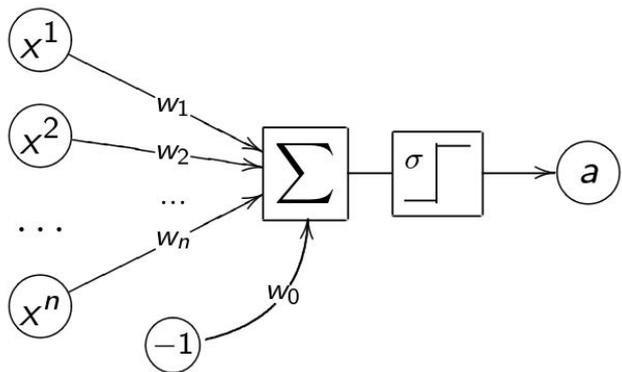
входной слой,
 n признаков

скрытый слой,
 H нейронов

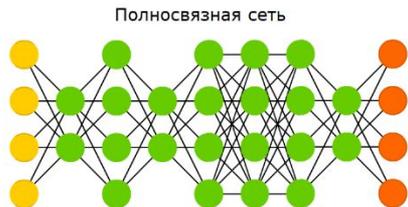
выходной слой,
 M нейронов



Модель нейрона.



$$a(x, w) = \sigma(\langle w, x \rangle) = \sigma\left(\sum_{j=1}^n w_j f_j(x) - w_0\right),$$



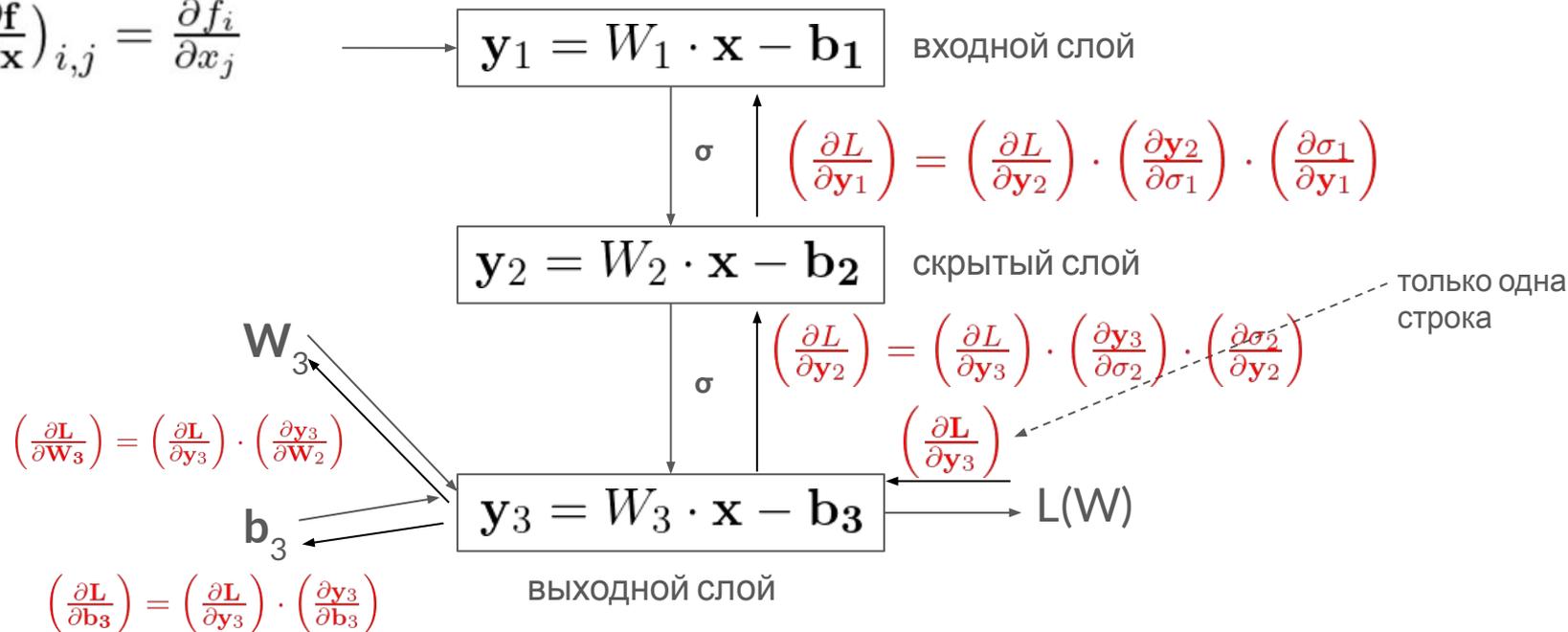
- Входной слой
- Скрытый слой
- Выходной слой

$$layer(\mathbf{x})_k = (\sigma \circ Linear)(\mathbf{x})_k = \sigma\left(-b_k + \sum_j W_{k,j} x_j\right)$$

Как от этого считать градиент по весам?

Обратное распространение ошибки

$$\left(\frac{\partial \mathbf{f}}{\partial \mathbf{x}}\right)_{i,j} = \frac{\partial f_i}{\partial x_j}$$



Vector-Jacobian (VJP) and Jacobian-Vector (JVP) products

$f : \mathbb{R}^n \rightarrow \mathbb{R}^m$ функция

$\partial f(x) \in \mathbb{R}^{m \times n}$ матрица Якоби

$(x, v) \mapsto \partial f(x)v$ JVP

$(x, v) \mapsto v\partial f(x),$ VJP

$\left(\frac{\partial L}{\partial \mathbf{y}_1}\right) = \left(\left(\frac{\partial L}{\partial \mathbf{y}_2}\right) \cdot \left(\frac{\partial \mathbf{y}_2}{\partial \sigma_1}\right)\right) \cdot \left(\frac{\partial \sigma_1}{\partial \mathbf{y}_1}\right)$ тоже VJP!

$$\partial(\mathbf{X}\mathbf{Y}) = (\partial\mathbf{X})\mathbf{Y} + \mathbf{X}(\partial\mathbf{Y}) \quad (37)$$

$\mathbf{v}\partial(W\mathbf{x} - \mathbf{b}) = \mathbf{v}W$ Сложность вычисления сравнима с прямым!

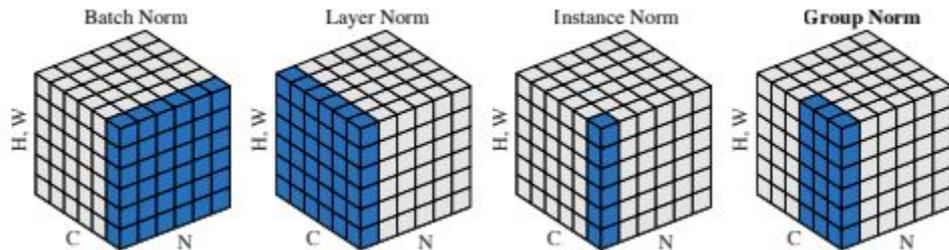
Нормализации пакетные, по слою, ...

$$\mu := \alpha\mu + (1 - \alpha)\mu_{\mathcal{B}},$$

$$\sigma^2 := \alpha\sigma^2 + (1 - \alpha)\sigma_{\mathcal{B}}^2,$$

$$\hat{x} = BN(x) = \gamma \frac{x - \mu}{\sqrt{\sigma^2 + \epsilon}} + \beta$$

Мотивация — хотим увеличить количество слоев, но градиенты экспоненциально растут или убывают на слоях



Сверточные слои

Input image

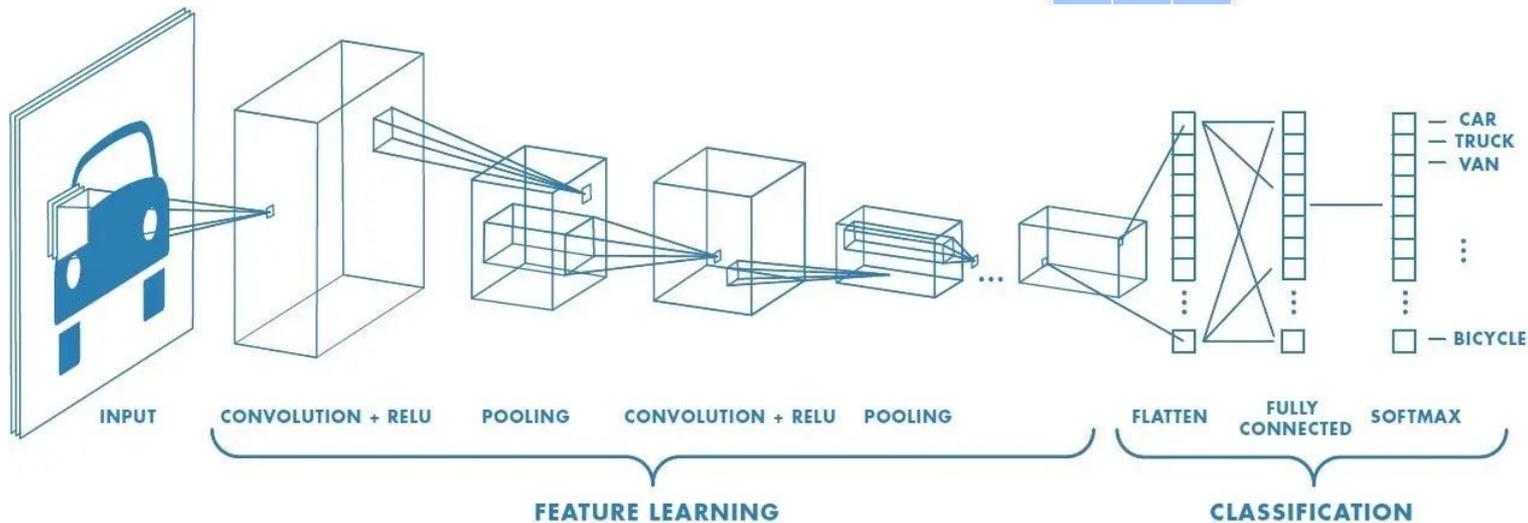
9	4	1	2	2
1	1	1	0	4
1	2	1	0	0
1	0	0	2	0
9	6	7	4	0

Filter

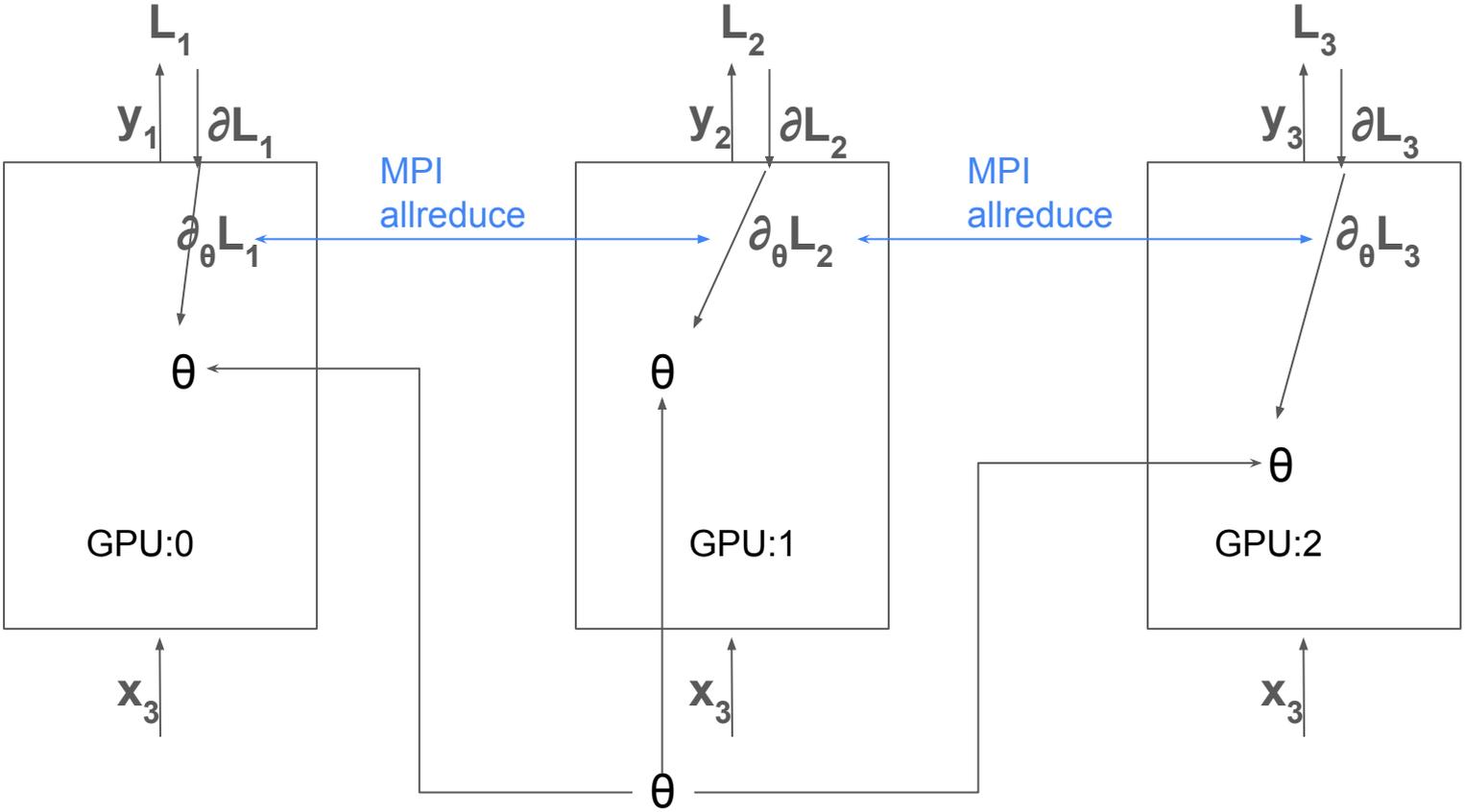
0	2	1
4	1	0
1	0	1

Output array

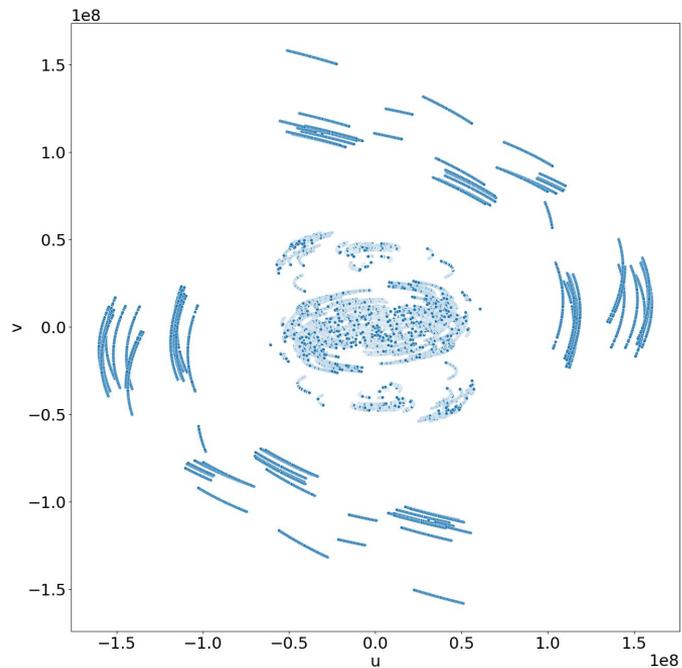
$$\begin{aligned} \text{Output}[0][0] &= (9*0) + (4*2) + (1*4) \\ &+ (1*1) + (1*0) + (1*1) + (2*0) + (1*1) \\ &= 0 + 8 + 1 + 4 + 1 + 0 + 1 + 0 + 1 \\ &= 16 \end{aligned}$$



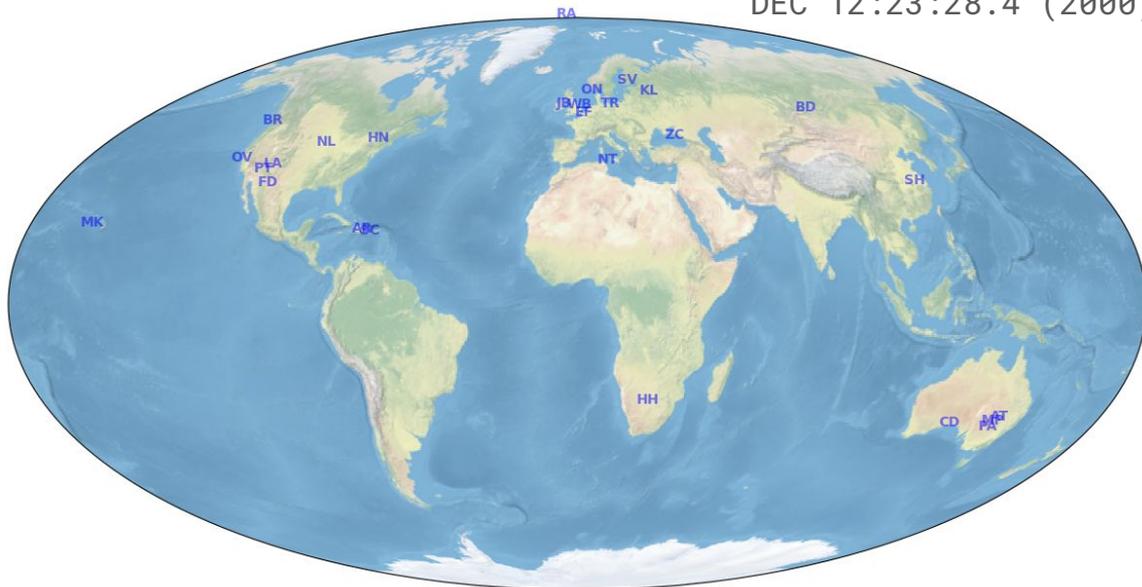
Distributed Data Parallel



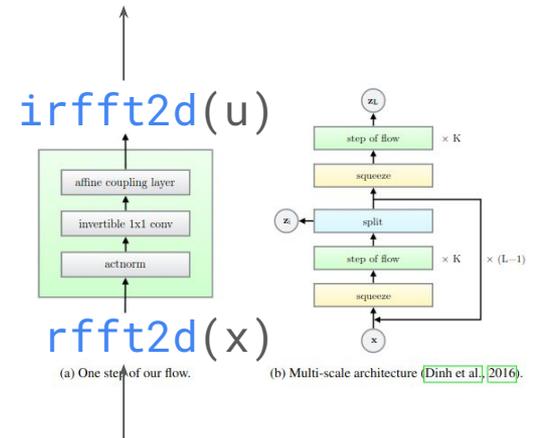
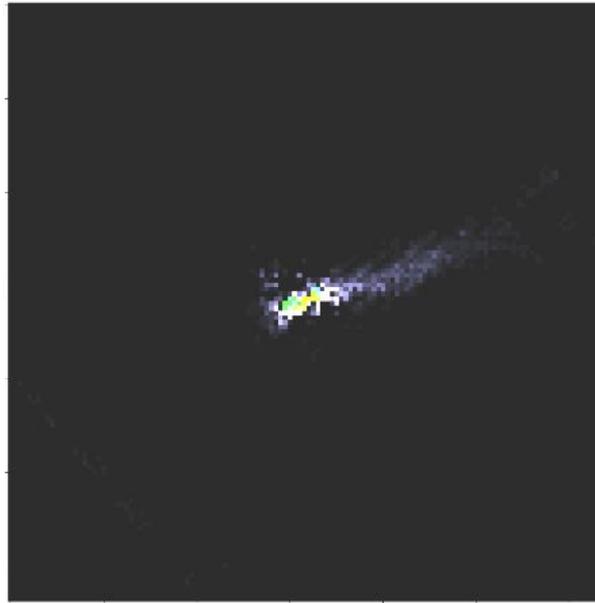
Данные наблюдений M87.



1228+126 LL-POL
1652.25 MHz
RA 12:30:49.4 (2000)
DEC 12:23:28.4 (2000)



Блок быстрого преобразования Фурье



Планы. Теорема об универсальной аппроксимации для потоковых нейросетей

Neural Autoregressive Flows

Chin-Wei Huang^{1,2*} David Krueger^{1,2*} Alexandre Lacoste² Aaron Courville

Abstract

Normalizing flows and autoregressive models have been successfully combined to produce state-of-the-art results in density estimation, via Masked Autoregressive Flows (MAF) (Papamakarios et al., 2017), and to accelerate state-of-the-art WaveNet-based speech synthesis to 20x faster than real-time (Oord et al., 2017), via Inverse Autoregressive Flows (IAF) (Kingma et al., 2016). We unify and generalize these approaches, replacing the (conditionally) affine univariate transformations of MAF/IAF with a more general class of invertible univariate transformations expressed as monotonic neural networks. We demonstrate that the proposed **neural autoregressive flows (NAF)** are universal approximators for continuous probability distributions, and their greater expressivity allows them to better capture multimodal target distributions. Experimentally, NAF yields state-of-the-art performance on a suite of density estimation tasks and outperforms IAF in variational autoencoders trained on binarized MNIST.¹

implicit losses, in the case of adversarial (Rezende & Mohamed, 2015), they can be used to approximate more complex distributions. This is a poor variational approximation to the target distribution to reflect the right amount of uncertainty (Turner & Sahani, 2011), resulting in unreliable predictions. We are thus interested in techniques for normalizing flows.

Recent work by Kingma et al. (2016) treats generative models as invertible transformations for constructing normalizing flows. The iterative process, unlike sampling from the autoencoder, and thus can be accelerated. This allows multiple layers to be stacked, increasing expressiveness for generative models (Papamakarios et al., 2016) or better for inference (Kingma et al., 2016) or better for generative models (Papamakarios et al., 2016). This also makes it possible to improve on the sequential conditional factorization assumed by autoregressive models such as PixelRNN or PixelCNN (Oord et al., 2016), and thus define a more flexible joint probability.

We note that the normalizing flow introduced by Kingma et al. (2016) only applies an affine transformation of each scalar random variable. Although this transformation is

Theorem 1. (DSF universally transforms any random variables into any desired random variables) *Let X be a random vector in an open set $\mathcal{U} \subset \mathbb{R}^m$. Let Y be a random vector in \mathbb{R}^m . Assume both X and Y have a positive and continuous probability density distribution. Then there exists a sequence of functions $(K_n)_{n \geq 1}$ parameterized by autoregressive neural networks in the following form*

$$K(\mathbf{x})_t = \sigma^{-1}(\mathcal{S}(x_t; \mathcal{C}_t(x_{1:t-1}))) \quad (14)$$

where $\mathcal{C}_t = (a_{tj}, b_{tj}, \tau_{tj})_{j=1}^n$ are functions of $x_{1:t-1}$, such that $Y_n \doteq K_n(X)$ converges in distribution to Y .

1. Introduction